

1 **Dynamics and acoustics of a spherical bubble**
2 **rising under gravity in an inviscid liquid**

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Abstract

12

13 The rising motion and the acoustic emission of a pulsating spherical gas/vapour bubble
14 in an isochoric, inviscid liquid are investigated. The motion is driven by the uniform and
15 constant force field due to the gravity. The liquid is assumed at rest at the initial time.
16 Unlike previous work on this subject, the mass of the bubble is not neglected, so that the
17 bubble motion is accurately simulated also in the presence of large volume variations. After
18 developing the relationships between the bubble motion to the liquid flow, a system of two
19 nonlinear ODEs for the radius and the position of the center of mass of the bubble is written.
20 The near-field pressure disturbance produced in the liquid by the bubble motion is evaluated
21 by means of elliptic integrals and an efficient approximation of it free from these special
22 functions is also used. The numerical integration of the ODE system allows to evaluate the
23 acoustic signal. This is carried out with the above mentioned approximation, and several
24 features of it are demonstrated through the study of a sample flow.

25 **I. INTRODUCTION**

26 The phenomenon of cavitation is, in many cases, a critical issue for the hydrodynamic
27 performance of the propulsion system of a ship, as well as for the thereby radiated noise.

28 Along with a large effort in the development of CFD codes, aimed at improving the
29 capability of simulating multiphase flows, significant work is devoted within the fluid
30 dynamics community to achieve more insight in cavitation through accurate modeling of
31 the relevant bubble dynamics (Brennen, 1995). Crucial points are the motion of the bubble
32 as a whole and the evolution of its volume in connection with the pressure field in the
33 surrounding liquid.

34 Over the last twenty years a number of papers were devoted to several aspects of the
35 evolution of gas (and vapour) bubbles moving within a liquid. In the following are
36 mentioned contributions aimed at improving understanding of the interaction between the
37 translating motion of a bubble and the simultaneous variation of its size.

38 A first group of papers deal with the coupled dynamics of a bubble moving under the
39 effect of pressure gradient due to a forcing acoustic pressure field. It is well represented by
40 papers of Reddy and Szeri (2002a,b), and Doinikov (2002, 2004). In all of the above papers
41 the radial motion of the bubble is determined through a form of the Rayleigh-Plesset
42 equation, somehow corrected to take into account the compressibility of liquid. In Reddy
43 and Szeri (2002a) a model derived by Magnaudet and Legendre (1998) is used in order to
44 write the equation for the translating motion coupled with that for radial dynamics of the
45 spherical bubble. In Reddy and Szeri (2002b) the same system of equations is perturbed,
46 following a method proposed by Plesset (1954), to study the stability of spherical shape

47 under the coupling with translational motion. In Doinikov (2002) both equations are
48 derived using a Lagrangian formulation, and then correction for liquid compressibility are
49 applied to the equation for radial motion in the form proposed by Keller and Miksis (1980).
50 In Doinikov (2004) the mutual interaction of volume pulsation, translational motion and
51 shape oscillation (the two latter assumed to be small) is studied, by expanding surface
52 modes in a series of Legendre polynomials, obtaining coupled equations to the second order
53 in the interaction terms. Shaw (2006) extends the analysis of Doinikov (2004), by including
54 third order interactions in the bubble axial velocity and shape modes, to demonstrate how
55 the interaction of shape modes can affect bubble motion and volume oscillation. In all
56 cases the ratio of the densities of fluid inside and outside the bubble does not play any role,
57 that is the density of the bubble is neglected.

58 A second group of works deals with the problem (not considered here) of the
59 interaction between shape oscillation and translating motion of a bubble rising due to the
60 gravitational force. Yang *et al.* (2003) investigate the rectilinear rise of a bubble
61 undergoing a transient behavior: the evolution of the bubble shape from release to steady
62 state is numerically simulated, while the ascent of a bubble which expands, or contracts,
63 due to a change in the ambient pressure is analytically studied. Klaseboer and Khoo
64 (2006) study the effect of the presence of solid walls near the bubble, along with the effect
65 of gravity. To do so they introduce equivalent bubble radius and wall velocity to extend the
66 solution of Rayleigh-Plesset equation to the case of a rising non-spherical bubble. Gordillo
67 *et al.* (2012) compared the results of their theoretical model of a rising bubble with the
68 shape deformation numerically simulated by means of a Navier-Stokes code.

69 Finally, in Tuteja *et al.* (2010) the motion of a spherically pulsating gas bubble under

70 gravity in an inviscid fluid is investigated, based on developments from previous work
71 (Chakraborty, 1990; Chakraborty and Tuteja, 1993), where the Lagrange formalism was
72 used to derive the equations of motion. The study is focused on the effects of changing
73 gravity on the bubble motion in a vertical plane. Tuteja *et al.* (2010) is the closest
74 reference to the work presented here: when the horizontal component of the velocity of the
75 center of mass is set to zero, the model reduces to one describing the vertical motion of the
76 bubble coupled with the spherical pulsation of its surface. The density of the bubble is
77 neglected in the comparison with that of the displaced liquid.

78 Significant effort has been spent to study the noise emissions from bubbles.
79 Outstanding contributions come from experimental investigation of specific problems in
80 practical applications as, for example, some works of Ceccio (Ceccio and Brennen, 1991;
81 Choi and Ceccio, 2007; Chang and Ceccio, 2011). Despite a deep interest in cavitation
82 noise, it is quite difficult to find experimental data for free-space far-field acoustics of a
83 single rising bubble. So, illustrating the features of the model presented here mostly rely on
84 theoretical arguments, for which a suitable reference is found in the conceptual
85 developments of Leighton (1994). Comparison with an experiment is proposed at the end,
86 where the acoustic far-field of the rising bubble has been calculated in conditions derived
87 from the data of a famous paper by Strasberg (1956).

88 In the present paper, dynamics and acoustics of a spherical gas/vapour bubble rising
89 in an isochoric, inviscid liquid are investigated. In Section II a system of two nonlinear
90 ODEs describing the dynamics of the spherical bubble is deduced, by relating the velocity
91 potential in the liquid with the bubble motion induced by the gravity field. A modified
92 form of the Rayleigh-Plesset equation describes the evolution in time of the bubble radius,

93 while the other equation gives the position of the center of mass during the rising motion.
94 Both equations are nonlinearly coupled and, unlike all papers referenced above, the ratio of
95 the densities inside and outside the bubble is taken into account as a parameter of the
96 model, and it is shown to be meaningful in certain conditions. In Section III, according to
97 a solution proposed by Farassat (2007) to the equation of Ffowcs Williams and Hawkings
98 (1969), the near-field pressure generated in the liquid is evaluated by means of elliptic
99 integrals and an efficient approximation of it, free from these special functions, is used
100 instead. It is based on a Maclaurin expansion in the ratio between the bubble radius and
101 the distance between observer and center of mass of the bubble. The bubble motion and
102 the related (approximation of) acoustic disturbance are numerically evaluated in
103 Section IV, where a comparison is made with the signature generated by a bubble formed
104 at the release of air from a nozzle in water, illustrated in a famous paper on the subject
105 (Strasberg, 1956). Some qualitative aspects of the dynamics and acoustics of the rising
106 bubble are illustrated in sample flows. Finally, conclusions are offered in Section V,
107 together with a brief description of the future work.

108 **II. EQUATIONS OF MOTION OF THE SPHERICAL BUBBLE**

109 The bubble, shown in Figure 1, is assumed spherical at any time and its center of
110 mass is thought of as superimposed to the center of the sphere. The radius (R) and the
111 position of the center of mass (\mathbf{x}_{cm}) of the bubble depend on time (t). They are solution of
112 a second order Cauchy differential problem that will be written below. Inside the bubble
113 there is a gas/vapour mixture which does not diffuse across the bubble boundary and is
114 subjected to isothermal compressions, $R(t) < R_0$, or expansions, $R(t) > R_0$, R_0 being the

115 radius at the initial time ($t = 0$). The surrounding liquid is homogeneous, inviscid and
 116 moves in an isochoric way, *i.e.* its density ρ is constant in space and time. It is assumed at
 117 rest at the initial time, therefore its motion remains irrotational at any successive time. On
 118 the fluids acts a constant and uniform gravitational field directed along the negative z -axis,
 119 g being the modulus of the corresponding acceleration.

120 A. Flow field around the bubble

121 The velocity potential in the liquid can be written in spherical coordinates by means
 122 of the well-known Legendre expansion:

$$\varphi(r, \mu; t) = \sum_{k=0}^{\infty} c_k(t) \frac{P_k(\mu)}{r^{k+1}} + \sum_{m=1}^{\infty} d_m(t) P_m(\mu) r^m. \quad (1)$$

123 In equation (1), μ is the cosine of the colatitude θ (along the z -axis, $\theta = 0$ for $z > z_{\text{cm}}$,
 124 $\theta = \pi$ for $z < z_{\text{cm}}$), P_l the l th Legendre polynomial and r is the distance from the origin.
 125 The potential is assumed as vanishing at infinity ($r \rightarrow \infty$), so that $d_m = 0$ for any m .

126 On the bubble boundary ($r = R$), the normal component of the velocity of the liquid
 127 (u_r) must take into account of the time derivative of the radius (\dot{R}) and the bubble
 128 translation velocity (\dot{z}_{cm}), by means of the Neumann condition:

$$u_r[R(t), \mu; t] = \frac{\partial \varphi}{\partial r}[R(t), \mu; t] = \dot{R}(t) P_0(\mu) + \dot{z}_{\text{cm}}(t) P_1(\mu).$$

129 It relates the above Legendre coefficients to the bubble motion in the following way:

$$c_0 = -R^2 \dot{R}, \quad c_1 = -\frac{1}{2} R^3 \dot{z}_{\text{cm}} \quad c_k = 0 \text{ for } k \geq 2. \quad (2)$$

130 Moreover, on the same boundary the pressure on the gas/vapour face (p_b) is balanced by
 131 the pressure p acting on the liquid face and by the surface tension: $p_b = p + 2\sigma/R$, where σ

132 is the surface tension coefficient of the gas-liquid interface. As a consequence, p_b is not
 133 uniform. It is expanded in the Legendre series:

$$\frac{p_b}{\rho}(\mu; t) = \frac{\bar{p}_b}{\rho}(t) + \sum_{k=1}^{\infty} \pi_k(t) P_k(\mu), \quad (3)$$

134 in which its mean value on the sphere \bar{p}_b is usually written as the sum of the constant
 135 vapour pressure p_v (at the ambient temperature) and the gas pressure p_g , related to the
 136 bubble volume V by the isothermal equation of state:

$$p_g(t)V(t) = p_g(0)V(0). \quad (4)$$

137 The remaining Legendre coefficients π_k ($k \geq 1$) have to be calculated, as discussed below.

138 Once the pressure-like quantity:

$$P(t) := \frac{\bar{p}_b(t) - p_\infty}{\rho} + gz_{\text{cm}}(t) - \frac{2\sigma}{\rho R(t)} \quad (5)$$

139 is introduced, the Legendre coefficients π_k for $k \geq 1$ (3) are related to the bubble motion by
 140 means of the Bernoulli law

$$-\frac{\partial\varphi}{\partial t} = \frac{|\mathbf{u}|^2}{2} + \frac{p - p_\infty}{\rho} + gz,$$

141 p_∞ being the asymptotic value assumed as $|\mathbf{x}| \rightarrow \infty$ by $p + \rho gz$. It follows:

$$\begin{aligned} -\frac{\dot{c}_0}{R} P_0 - \frac{c_0 \dot{z}_{\text{cm}} + \dot{c}_1}{R^2} P_1 - \frac{2c_1 \dot{z}_{\text{cm}}}{R^3} P_2 = \\ = \left(\frac{1}{2} \frac{c_0^2}{R^4} + \frac{c_1^2}{R^6} + P \right) P_0 + \left(2 \frac{c_0 c_1}{R^5} + gR + \pi_1 \right) P_1 + \left(\frac{c_1^2}{R^6} + \pi_2 \right) P_2 + \sum_{k=3}^{\infty} \pi_k P_k. \end{aligned} \quad (6)$$

142 B. Dynamics of radius and of center of mass of the bubble

143 The first consequence of relation (6) is that $\pi_k = 0$ for $k \geq 3$. Moreover, its projection
 144 along P_0 leads to the time evolution equation of the bubble radius:

$$\ddot{R} = \frac{1}{R} \left(-\frac{3}{2} \dot{R}^2 + \frac{1}{4} \dot{z}_{\text{cm}}^2 + P \right). \quad (7)$$

145 The differential equation (7) being of the second order, the initial values of R , *i.e.*

146 $R(0) := R_0$, and of its time derivative, *i.e.* $\dot{R}(0) := \dot{R}_0$, have to be prescribed.

147 The position of the center of mass of the bubble, \mathbf{x}_{cm} , is written in terms of the
148 density ρ_b of the gas/vapour mixture as:

$$\mathbf{x}_{\text{cm}}(t) = \frac{1}{m_b} \int_{V(t)} dV(\mathbf{x}; t) \rho_b(\mathbf{x}; t) \mathbf{x},$$

149 $m_b = \bar{\rho}_b V$ being the constant mass of gas/vapour trapped inside the bubble. As a

150 consequence, the mean density of the mixture changes according to the following equation:

$$\bar{\rho}_b(t)V(t) = \bar{\rho}_b(0)V(0). \quad (8)$$

151 The mixture moves according to the equation:

$$\rho_b D_t \mathbf{u} = -\nabla p_b - \rho_b g \mathbf{e}_z$$

152 and then the second time derivative of the position of the center of mass can be written as:

$$\ddot{\mathbf{x}}_{\text{cm}} = \frac{1}{m_b} \int_V dV \rho_b D_t \mathbf{u} = -\frac{1}{m_b} \int_{\partial V} dS p_b \mathbf{n} - g \mathbf{e}_z. \quad (9)$$

153 The only nonvanishing component of the right-hand side is along the z -axis, so that if the

154 initial position and velocity of the center of mass are prescribed along z , it turns out to be

155 $\mathbf{x}_{\text{cm}} \equiv (0, 0, z_{\text{cm}})^T$ at any successive time.

156 By using the balance of the normal stresses at the bubble interface and the Bernoulli

157 law in the liquid, p_b is written as $p + 2\sigma/R$, that is

$$p_b = p_\infty - \rho g z - \rho \left(\frac{\partial \varphi}{\partial t} + \frac{|\mathbf{u}|^2}{2} \right) + 2 \frac{\sigma}{R},$$

158 so that the integral (9) can be evaluated. Note that it includes \dot{c}_1 which, in turn, depends

159 on the acceleration \ddot{z}_{cm} , as shown by the second relation (2). The resulting equation is

160 solved with respect to \ddot{z}_{cm} and, once the ratio between the mean gas/vapour and the liquid
 161 densities:

$$\varepsilon(t) := \frac{\bar{\rho}_{\text{b}}(t)}{\rho} \quad (10)$$

162 has been introduced, the equation of motion of the center of mass of the bubble follows:

$$\ddot{z}_{\text{cm}} = 2g \frac{1 - \varepsilon}{1 + 2\varepsilon} - \frac{3}{1 + 2\varepsilon} \frac{\dot{R}}{R} \dot{z}_{\text{cm}}. \quad (11)$$

163 This equation reduces to the one for the dynamics of a rigid sphere (see Batchelor, 1967,
 164 Equation 6.8.20) by enforcing $R \equiv \text{constant}$, and to the equation for vertical motion in
 165 Tuteja *et al.* (2010) when taking $\varepsilon = 0$. As usual, the evolution equation (9) being of the
 166 second order, the position, *i.e.* $z_{\text{cm}}(0) =: z_{\text{cm}0}$, and the velocity, *i.e.* $\dot{z}_{\text{cm}}(0) =: \dot{z}_{\text{cm}0}$, of the
 167 center of mass at the initial time have to be assigned.

168 Finally, from the Bernoulli law (6) and from the dynamics of the center of mass (11)
 169 the two non-vanishing Legendre coefficients of the pressure p_{b} on the vapour/gas face of
 170 the bubble interface are obtained:

$$\pi_1 = \frac{3\varepsilon}{1 + 2\varepsilon} (\dot{R}\dot{z}_{\text{cm}} - gR), \quad \pi_2 = \frac{3}{4} \dot{z}_{\text{cm}}^2. \quad (12)$$

171 On the basis of the above considerations, the motion of the bubble is defined by the
 172 solution of the nonlinear ODE system formed by the evolution equations of the radius (7)
 173 and of the position of the center of mass (11), together with the isothermal equation of
 174 state (4), the mass conservation (8) and the definitions of P (5) and of ε (10). In the
 175 present paper, this system is numerically integrated by means of a fourth order
 176 Runge-Kutta method. In the following section, the evaluation of the pressure disturbance
 177 due to the bubble motion is briefly described.

178 **III. ACOUSTICS OF THE RISING SPHERICAL BUBBLE**

179 The pressure perturbation $\mathcal{P}(\mathbf{x}_o; t)$ radiated from the rising bubble and received, at
 180 time t , by an observer placed at the point $\mathbf{x}_o = (r_o \cos \beta_o, r_o \sin \beta_o, z_o)^T$ outside the bubble
 181 (see Figure 1) is now evaluated. It results to be a function of the distance of the observer
 182 from the z -axis (r_o), of the difference $z_o - z_{cm} =: \zeta$ and of the time.

183 The distance of the observer position from any given point

184 $\mathbf{x} = (R \sin \theta \cos \beta, R \sin \theta \sin \beta, z_{cm} + R \cos \theta)^T$ of the bubble surface is:

$$|\mathbf{x}_o - \mathbf{x}| = (R^2 - 2R\zeta \cos \theta + \zeta^2 + r_o^2)^{1/2} [1 - \chi^2 \cos(\beta - \beta_o)]^{1/2}, \quad (13)$$

185 with $\chi \in [0, 1)$ given by

$$\chi^2 = \frac{2Rr_o \sin \theta}{R^2 - 2R\zeta \cos \theta + \zeta^2 + r_o^2}.$$

186 Note that $\chi \rightarrow 1$ as $r_o \rightarrow R \sin \theta$ and $\zeta \rightarrow R \cos \theta$, *i.e.* for \mathbf{x}_o approaching the bubble
 187 surface.

188 The expression of the acoustic pressure disturbance can be written according to

189 Ffowcs Williams and Hawkings (1969), neglecting the contribution of volume sources, that

190 is including only the so-called *thickness* and *loading* terms. However, the form adopted

191 here is a solution of the Ffowcs Williams and Hawkings equation derived by Farassat

192 (2007), known as *formulation 1*. Along with symbols defined so far, one further introduces:

193 $\boldsymbol{\xi} = \mathbf{x}_o - \mathbf{x}$, $\xi = |\boldsymbol{\xi}|$, $u_\xi = \mathbf{u} \cdot \boldsymbol{\xi} / \xi$ (the component of bubble velocity along $\boldsymbol{\xi}$) and the source

194 time, t_s . Farassat's formulation 1 can be written as:

$$\begin{aligned}
 \mathcal{P}(\mathbf{x}_o; t) &= \frac{\rho}{4\pi} \frac{\partial}{\partial t} \int_{\partial V} dS \left[\frac{u_n}{\xi(1 - u_\xi/c_\infty)} \right]_{t_s=t_e} \\
 &+ \frac{1}{4\pi c_\infty} \frac{\partial}{\partial t} \int_{\partial V} dS \left[\frac{p \mathbf{n} \cdot \boldsymbol{\xi}}{\xi^2(1 - u_\xi/c_\infty)} \right]_{t_s=t_e} \\
 &+ \frac{1}{4\pi} \int_{\partial V} dS \left[\frac{p \mathbf{n} \cdot \boldsymbol{\xi}}{\xi^3(1 - u_\xi/c_\infty)} \right]_{t_s=t_e}, \tag{14}
 \end{aligned}$$

195 where c_∞ is the speed of sound in the undisturbed medium. Under the integral sign, \mathbf{u} , \mathbf{n}
 196 (then u_n) and p are to be considered as functions of \mathbf{x} and t_s . Integration is performed on
 197 the surface of the bubble. The subscript $t_s = t_e$ means that values of integrands, at each
 198 point of the surface, are calculated at the proper *emission time*, determined through the
 199 solution of:

$$|\mathbf{x}_o - \mathbf{x}(t_e)| = c_\infty(t - t_e).$$

200 There is evidence that, in most cases, far-field acoustic pressure from bubbles or
 201 cavities is due to rapid oscillations of their volume. This is represented by the first term in
 202 equation (14). However, in order to have a clear picture of all the mechanisms at work in
 203 generating pressure disturbances from a moving bubble, it is instructive to look at the
 204 (very) near-field behaviour of the system. Here the wave equation is dominated by the
 205 Laplacian (Lezzi and Prosperetti, 1986): in fact, by letting $c_\infty \rightarrow \infty$ in equation (14), one
 206 obtains

$$\mathcal{P}(\mathbf{x}_o; t) = \frac{\rho}{4\pi} \frac{\partial}{\partial t} \int_{\partial V} dS(\mathbf{x}; t_s) \frac{u_n(\mathbf{x}; t_s)}{|\mathbf{x}_s - \mathbf{x}|} + \frac{1}{4\pi} \int_{\partial V} dS(\mathbf{x}; t_s) \tilde{p}(\mathbf{x}; t_s) \mathbf{n}(\mathbf{x}; t_s) \cdot \frac{\mathbf{x}_o - \mathbf{x}}{|\mathbf{x}_o - \mathbf{x}|^3}, \tag{15}$$

207 This means neglecting the time due to propagation of sound from the source to the
 208 observer, thus considering the system as *instantaneous*, rather than *retarded*. It seems

209 reasonable, thinking of the large speed of sound in water and the small source-observer
 210 distance for the case examined. In equation (15) the normal velocity u_n and the pressure \tilde{p}
 211 at a point $\mathbf{x} \in \partial V$ are as follows

$$u_n = \dot{R} + \dot{z}_{\text{cm}} \cos \theta, \quad \tilde{p} = \pi_1 \cos \theta + \frac{3}{2} \pi_2 \cos^2 \theta,$$

212 where π_1 and π_2 are given in equations (12). In fact, \tilde{p} is the part of p , on the bubble
 213 surface, exceeding the uniform term

$$\frac{\bar{p}_b}{\rho} - \frac{2\sigma}{\rho R} - \frac{\pi_2}{2},$$

214 which is going to give a vanishing contribution to the relevant integral in equation (15).

215 Using the distance (13) the pressure disturbance is evaluated in terms of the complete
 216 elliptic integrals of the first kind K and of the second kind E as:

$$\begin{aligned} \mathcal{P}(\mathbf{x}_o; t) = & \frac{\rho}{\pi} \frac{\partial}{\partial t} \left[R^2 \int_0^\pi d\theta \sin \theta \frac{\dot{R} + \dot{z}_{\text{cm}} \cos \theta}{D} K(k) \right] \\ & + \frac{\rho R^2}{4\pi} \int_0^\pi d\theta \sin \theta \frac{\tilde{p}}{D} \left\{ \frac{4(\zeta \cos \theta - R)}{D^2} \frac{E(k)}{k'^2} + \frac{1}{R} \left[\frac{2 - k^2}{k'^2} E(k) - 2K(k) \right] \right\}. \end{aligned} \quad (16)$$

217 In equation (16) the quantity

$$D := [(r_o + R \sin \theta)^2 + (\zeta - R \cos \theta)^2]^{1/2}$$

218 has been introduced for short. Moreover, the modulus of the elliptic integral is:

$$k = \frac{2\sqrt{Rr_o \sin \theta}}{D}. \quad (17)$$

219 and the complementary modulus is $k' = (1 - k^2)^{1/2}$.

220 In the general case in which the observer does not lie on the z -axis ($r_o > 0$), \mathcal{P} is
 221 evaluated by deriving in time inside the integral (16) and then performing a numerical

222 integration. This latter is not needed for $r_o = 0$ (observer lying on the z -axis), that implies
 223 $k = 0$ by means of the definition of the modulus (17). Due to the fact that $K(0) = \pi/2$, the
 224 pressure perturbation \mathcal{P} (16) takes the form:

$$\mathcal{P}(\mathbf{x}_o; t) = \frac{\rho}{3} \frac{R^3}{|\zeta|} \left(3 \frac{\ddot{R}}{R} + 6 \frac{\dot{R}^2}{R^2} + \frac{6}{\zeta} \frac{\dot{R}}{R} \dot{z}_{\text{cm}} + \frac{1}{\zeta} \ddot{z}_{\text{cm}} + \frac{2}{\zeta^2} \dot{z}_{\text{cm}}^2 + \frac{1}{\zeta} \frac{\pi_1}{R} + \frac{6}{5\zeta^2} \pi_2 \right). \quad (18)$$

225 In the following results, the integral (16) is not used to calculate the pressure
 226 perturbation from the rising bubble. In fact, it can be successfully replaced by an
 227 approximation (named as \mathcal{P}^a), that is obtained by neglecting powers of order greater than
 228 3 of the ratio $\eta = R/\delta_o$, δ_o being the distance of the observer from the center of mass of the
 229 bubble, *i.e.*, $\delta_o = (r_o^2 + \zeta^2)^{1/2}$. Using appropriate Maclaurin expansions for powers of:

$$\frac{1}{|\mathbf{x}_o - \mathbf{x}|} = \frac{1}{\delta_o} \left\{ 1 - 2 \frac{R}{\delta_o} \left[\frac{\zeta}{\delta_o} \cos \theta + \frac{r_o}{\delta_o} \sin \theta \cos(\beta - \beta_o) \right] + \frac{R^2}{\delta_o^2} \right\}^{1/2} := \frac{1}{\delta_o} (1 - 2\eta\gamma + \eta^2),$$

230 in the parameter $-2\eta\gamma + \eta^2$, the pressure perturbation (15) is approximated by:

$$\begin{aligned}
 \mathcal{P}^a(\mathbf{x}_o; t) = \frac{\rho}{3} \frac{R^3}{\delta_o} \left[3 \frac{\ddot{R}}{R} + 6 \frac{\dot{R}^2}{R^2} + \frac{6\zeta}{\delta_o^2} \frac{\dot{R}}{R} \dot{z}_{\text{cm}} + \frac{\zeta}{\delta_o^2} \ddot{z}_{\text{cm}} + \frac{1}{\delta_o^2} \left(\frac{3\zeta^2}{\delta_o^2} - 1 \right) \dot{z}_{\text{cm}}^2 \right. \\
 \left. + \frac{\zeta}{\delta_o^2} \frac{\pi_1}{R} + \frac{3}{5\delta_o^2} \left(\frac{3\zeta^2}{\delta_o^2} - 1 \right) \pi_2 \right]. \quad (19)
 \end{aligned}$$

231 On the symmetry axis ($r_o = 0$, $\delta_o = |\zeta|$) the approximation (19) gives the exact
 232 representation (18).

233 The first five terms of equation (19) are generated by the first integral in the
 234 right-hand side of equation (15). This would give rise in the far field to the so-called
 235 *thickness noise*, and is related to the geometry and kinematic properties of the moving
 236 source. The first three terms, containing time derivatives of bubble radius, are responsible
 237 for the *monopole* character of the far-field noise of the bubble. The fourth and fifth terms

238 of equation (19), contain only derivatives of position of the center of mass of the bubble.
239 They clearly exhibit a different directivity pattern and a higher order dependence on the
240 inverse of the source-observer distance, which gives them a dipole character. In fact, they
241 share the structure of the last two terms of equation (19), generated by the second integral
242 in the right-hand side of equation (15). These, coming from bubble surface pressure
243 loading, are widely known as *dipole* terms. As it will be seen from the illustration of some
244 sample results, they give a smaller contribution compared to the corresponding terms from
245 thickness, affecting the absolute value of the peak disturbance, while not changing the
246 overall pattern.

247 Moving to the far field, the complete right-hand side of equation (14) should be
248 considered. However, it is clear from the previous analysis the growing dominance of
249 thickness terms from bubble radius (that is, volume) oscillation. As far as the second
250 integral in the right-hand side of equation (14) is concerned, one may observe that it is still
251 affected by the dependence on c_∞^{-1} . This results in a limited contribution to the overall
252 signal, compared to that generated by the first integral, at least in the generality of cases.
253 Particular situations, where it may be comparable to (or even dominant on) thickness
254 noise, may be caused by different patterns of bubble dynamics. One such case may be
255 given when rapid oscillations of surface pressure load are generated independently of the
256 radius pulsation, as in the interaction of bubble with unsteady vortex motion. A further
257 example is that provided in the case of phenomena causing intense shock waves to be
258 released by the collapse of a bubble in the vicinity of a solid surface, as it was shown in a
259 previous application of Ffowcs Williams and Hawkings formulation to bubble acoustics
260 (Jamaluddin *et al.*, 2011).

IV. NUMERICAL RESULTS

All the quantities are made nondimensional by adopting the initial bubble radius R_0 as scale of lengths, ρR_0^3 as scale of masses and the quantity $T_0 = 2\pi/\omega_0$ as scale of times, where ω_0 is the natural (circular) frequency of oscillation of the bubble, given as ω_N by Brennen (1995). There, the definition of ω_N is based on the *equilibrium radius* (the value of R for a bubble with a given amount of gas to be in equilibrium at the given condition of external pressure), while ω_0 is calculated here using R_0 , which is made the equilibrium radius at the given external pressure by properly choosing the amount of gas contained in the bubble. In the present calculations the bubble contains air at temperature 20 °C, moves in water by starting with radius $R_0 = 2$ mm and center of mass in $z_{\text{cm}} = -0.1$ m, so that T_0 results to be less than 1 ms. The surface tension coefficient σ at the bubble wall is 0.0729 N m⁻¹. In the following discussion, the symbols introduced so far to denote the dimensional quantities will be used for the corresponding nondimensional quantities, while, for the sake of clarity, initial data will be presented in dimensional form.

The main feature of the proposed model for the motion of the center of mass of the bubble and its consequences on the pressure perturbation are illustrated in Figures 2 and 3.

Neglecting gas density inside the bubble, that is enforcing $\varepsilon = 0$, appears to be a common practice in most previous papers referenced here (in particular, in Tuteja *et al.*, 2010), which use the Lagrangian formulation to derive the equations of motion. This causes significant differences in the peak values of \dot{z}_{cm} (Fig. 2(a)), and results in an overestimation of the vertical position of the center of mass at a given time (Fig. 2(b)). Although this difference may be considered as a negligible one, the overall effect on the

283 pressure perturbation is surprisingly significant. At the observer position considered here,
 284 the peak amplitudes exhibit a difference of around 20% (Fig. 3), and also a slight phase
 285 shift between two corresponding peaks appears (magnified in the zoomed view).

286 Next issue is looking at how the model reproduces the main features of bubble
 287 dynamics. Figure 4 shows the starting evolution of a bubble with initial velocities
 288 $\dot{R}_0 = 5 \text{ m s}^{-1}$ and $\dot{z}_{\text{cm}0} = 0 \text{ m s}^{-1}$. Typically, at the strong collapses of R (a) correspond
 289 large spikes in \dot{z} (b), causing the stepped trend of z (c), characteristic of the bubble rising
 290 motion.

291 In order to take a global look at the system behaviour, a case with a different
 292 combination of initial data is studied next. Fig. 5 shows radial and vertical motions of a
 293 bubble starting with zero radial velocity and nonzero vertical velocity. These are compared
 294 with the analogous motions consequent to zero radial and zero vertical initial velocities, *i.e.*
 295 driven only by the effect of gravity. The latter curves are very smooth: they represent the
 296 rising motion, due to gravity, of a sphere steadily expanding because of the decreasing
 297 pressure along the rising path. The initial nonzero vertical velocity induces in the curve of
 298 the corresponding radial motion, shown in Fig. 5(a), an oscillating pattern around a global
 299 trend representing the steady expansion already observed in the curve of the motion with
 300 $\dot{z}_{\text{cm}0} = 0 \text{ m s}^{-1}$. The oscillation in this case is quite regular, and does not exhibit the
 301 violent collapse appearing in the pattern of Fig. 4, so that the curve in Fig. 5(b), showing
 302 the vertical motion of the center of mass, is not stepped: indeed it appears as smooth as
 303 the one of the vertical motion with $\dot{z}_{\text{cm}0} = 0 \text{ m s}^{-1}$.

304 Fig. 6 shows the pressure disturbance generated by the motion of the bubble at an
 305 observer position whose minimum distance from the moving source is 5 times the initial

306 bubble radius R_0 . The signature generated by the motion with $\dot{z}_{\text{cm}0} = 0 \text{ m s}^{-1}$ is smooth
 307 and closely resembles the acoustic time signature of a dipole source, like the one expected
 308 as the effect of a rigid body having a relative motion with respect to the observer. This is
 309 basically represented by the last two terms in equation (19), those depending on \ddot{z}_{cm} and
 310 \dot{z}_{cm}^2 . On the other hand, the signature generated by the motion with nonzero initial vertical
 311 velocity is clearly characterized by the oscillating pattern of the corresponding radial
 312 motion: it exhibits a significant monopole component, represented in equation (19) by
 313 terms containing time derivatives of R . This is superimposed to a dipole pattern similar to
 314 that of the curve with zero initial vertical velocity, although hidden by the large oscillations
 315 amplitude.

316 To better understand the combination of different effects in the resulting acoustic
 317 disturbance of the bubble, the signatures are filtered in time by calculating the average,
 318 $\mathcal{P}_{\text{ave}}^{\text{a}}$, and the root mean square, $\mathcal{P}_{\text{rms}}^{\text{a}}$ (with respect to $\mathcal{P}_{\text{ave}}^{\text{a}}$), of the function \mathcal{P}^{a} (De
 319 Bernardis and Riccardi, 2015). Thus, one finally gets:

$$\mathcal{P}_{\text{ave}}^{\text{a}}(t) := \frac{1}{T} \int_{-T/2}^{T/2} d\tau \mathcal{P}^{\text{a}}(t + \tau),$$

$$\mathcal{P}_{\text{rms}}^{\text{a}}(t) := \sqrt{\frac{1}{T} \int_{-T/2}^{T/2} d\tau [\mathcal{P}^{\text{a}}(t + \tau) - \mathcal{P}_{\text{ave}}^{\text{a}}(t)]^2}$$

320 (the dependencies on \mathbf{x}_o have been omitted, for the sake of shortness). Here T is, in
 321 principle, an arbitrary time interval. In practice, it is chosen as the time interval
 322 corresponding to an integer number of complete oscillations. This results in the curves
 323 drawn in Figs. 7 and 8. Fig. 7 shows the average curves of the pressure signatures pictured
 324 in Fig. 6. The curve with $\dot{z}_{\text{cm}0} = 0 \text{ m s}^{-1}$ almost reproduces the original pressure

325 perturbation, saying that the averaging process is, more or less, uneffective in this case. On
 326 the contrary, the curve with $\dot{z}_{\text{cm}0} = 0.5 \text{ m s}^{-1}$ looks now very similar to the other one. The
 327 averaging process has removed the oscillating part of the signal, revealing the dipole-like
 328 carrier signature. Fig. 8 shows the root mean square curves of the pressure perturbation in
 329 Fig. 6. They are associated to the amplitude of the oscillations that have to be
 330 superimposed to the average curves in Fig. 7 in order to give back the pressure perturbation
 331 of Fig. 6. Finally, only for the case with $\dot{z}_{\text{cm}0} = 0.5 \text{ m s}^{-1}$, the pressure perturbation and
 332 the corresponding average and root mean square curves are reported in Fig. 9.

333 Equation (19) exhibits significant features. In fact, it describes the pressure
 334 disturbance generated by the bubble through the sum of several terms, each representing
 335 the role played in this phenomenon by a particular quantity related to the bubble
 336 dynamics. In particular, the first three terms, all including derivatives of R , represent the
 337 energy associated to the bubble pulsation, clearly exhibiting a monopole character. The
 338 last four terms (two including only derivatives of z_{cm} , two including components of the
 339 bubble surface pressure), show the dipole nature of the perturbation generated by a rigid
 340 body through its relative motion with respect to a steady observer. These two main
 341 components of equation (19) are separately drawn in Fig. 10, for the pressure perturbations
 342 of Fig. 9. In particular, one can observe that the second (dipole) part is very close to the
 343 average curve of the overall signature, as it is shown in Fig. 11. Here the dipole
 344 contributions from *thickness* (terms with derivatives of z_{cm}) and *loading* (terms with
 345 components of pressure) are plotted separately to show the relative magnitude.

346 Finally, Fig. 12 shows the result obtained in the calculation of the far-field acoustic
 347 signature of a rising bubble. It refers to the only (to authors' knowledge) available

348 experimental result for a single rising bubble in free space, so that it makes sense to have a
349 true far-field prediction.

350 The experiment, reported in Strasberg (1956), performs measurement of sound
351 generated by an air bubble released at a nozzle in water, next freely rising to gravity. The
352 reference picture (not reproduced here) shows, along with the measured pressure, the shape
353 of the rising bubble corresponding to selected times. Proper initial conditions for the
354 integration procedure are derived from the experimental set-up, following suggestions given
355 in the paper.

356 Because of the large distance of the observer, this comparison is not particularly
357 suitable to highlight the features of the formulation presented here. In fact, in this case the
358 acoustic signature could have been calculated considering just the *monopole thickness* (first
359 three) terms in equation (19), using the correction to properly include retarded times, after
360 correcting for compressibility effect also the dynamic equations. The remaining four
361 (*dipole*) terms provide negligible contribution.

362 However, the comparison would show a reasonable agreement with the measured
363 signature, with a 18% underestimation of the measured peak sound pressure (in the paper,
364 the calculated 40% overestimation is claimed as “... *in fair agreement* ...”!), and a slower
365 decay in the amplitude of the calculated waveform compared to the data. The departure
366 from the measured peak pressure may be partly related to the uncertainty in extracting
367 data from the picture of the measurement. If one looks at the picture, showing the
368 evolution of the bubble shape, it is a reasonable hypothesis to relate the underestimation of
369 the waveform decay to the constraint of spherical shape enforced in the mathematical
370 model.

371 V. CONCLUDING REMARKS AND PERSPECTIVES

372 Main purpose of this work is the qualitative investigation of the acoustic disturbance
 373 radiated by a spherical bubble rising under the effect of the gravity through an inviscid
 374 isochoric fluid. A modified Rayleigh-Plesset (RP) equation for the bubble radius has been
 375 introduced, together with an equation for the dynamics of the center of mass of the bubble.
 376 These equations exhibit a twofold nonlinear coupling. The equation for the radial motion,
 377 compared to the original form of the RP equation for a pulsating bubble at rest in an
 378 inviscid fluid, includes an additional term that is proportional to the kinetic energy of the
 379 translating motion. Equation (11) describing the motion of the center of mass is a general
 380 representation of a translating bubble with time-varying radius and is coupled to the RP
 381 equation through a mixed term containing the product $\dot{R}\dot{z}_{\text{cm}}$. It includes, as particular
 382 cases, the equation for the vertical motion of a rigid sphere ($R \equiv \text{constant}$) (Batchelor,
 383 1967), as well as the equation for the rising motion of a bubble proposed by Tuteja *et al.*
 384 (2010).

385 It is worth noticing that in the extensive literature summarized in the Introduction,
 386 mostly based on the Lagrangian formulation for the derivation of the equations of motion,
 387 it is a common practice to neglect the ratio ε between the density of the mixture inside the
 388 bubble and the one of the external liquid. It has been shown here that it may significantly
 389 affect the pressure disturbance, in particular when the bubble motion leads to large volume
 390 variations.

391 A further remark concerns the closed form (19), proposed for estimating the acoustic
 392 signature radiated by the rising bubble. It was already observed that this formula allows

393 for the qualitative understanding of the roles played by velocities and accelerations in the
394 acoustic emission of the rising bubble. One more feature has been actually shown here:
395 formula (19) separates monopole and dipole contributions, and these components appear to
396 be closely related to the root mean square and average, respectively, of the acoustic
397 signature. Although the formula was derived to propose a qualitative study of the bubble
398 pressure disturbance in the near field, it has been successfully used, properly adapted along
399 with the dynamic equations, to simulate the pressure signature measured in a laboratory
400 experiment.

401 The results of the numerical integration of the system of equations for the dynamics of
402 the rising bubble have been successfully compared with those of a three-dimensional
403 boundary element code, which simulates the potential flow in the liquid. In the near
404 future, this code will be used for investigating the departures of the bubble shape from the
405 spherical one during the rising, as well as for estimating their effects on the acoustics.
406 Furthermore, it will be used to study the dynamics, and thereby generated sound, of a
407 non-spherical rising bubble. Analytical and numerical extensions of the present approach
408 to the bubble rising in presence of a wall or of a (fixed) free surface are also under
409 investigation at the present time.

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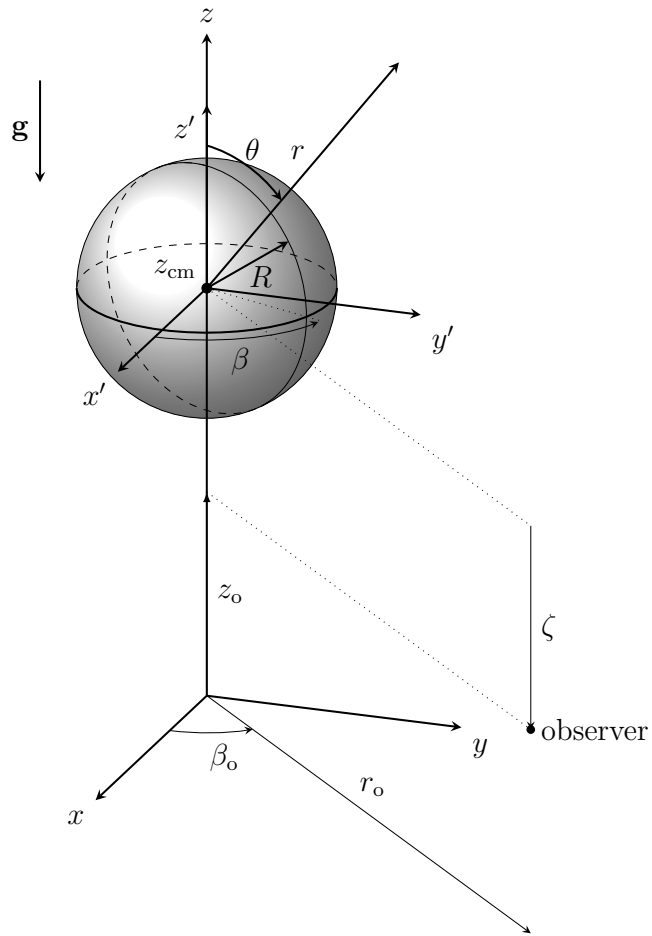
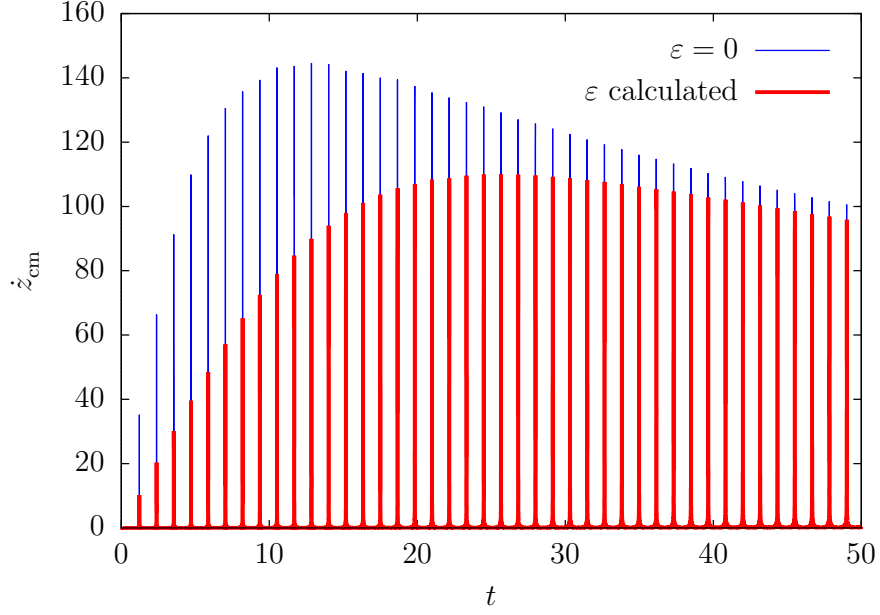
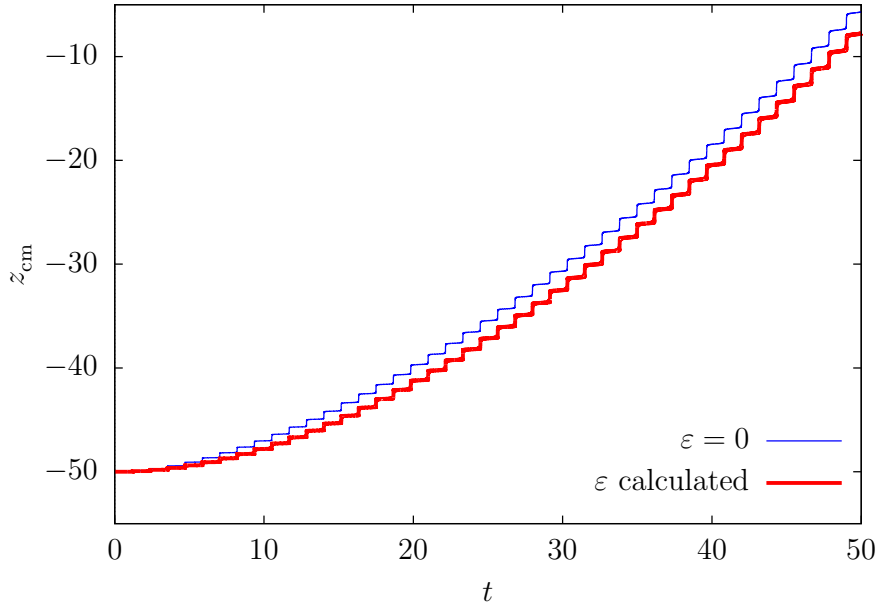


Figure 1: Schematic for bubble dynamics and acoustics.



(a)



(b)

Figure 2: (Color online) Velocity (a) and position of the center of mass (b) of the bubble with density ratio ε calculated or enforced as $\varepsilon = 0$. Initial data: $\dot{R}_0 = 20 \text{ m s}^{-1}$, $z_{\text{cm}0} = -0.1 \text{ m}$, $\dot{z}_{\text{cm}0} = 0 \text{ m s}^{-1}$.

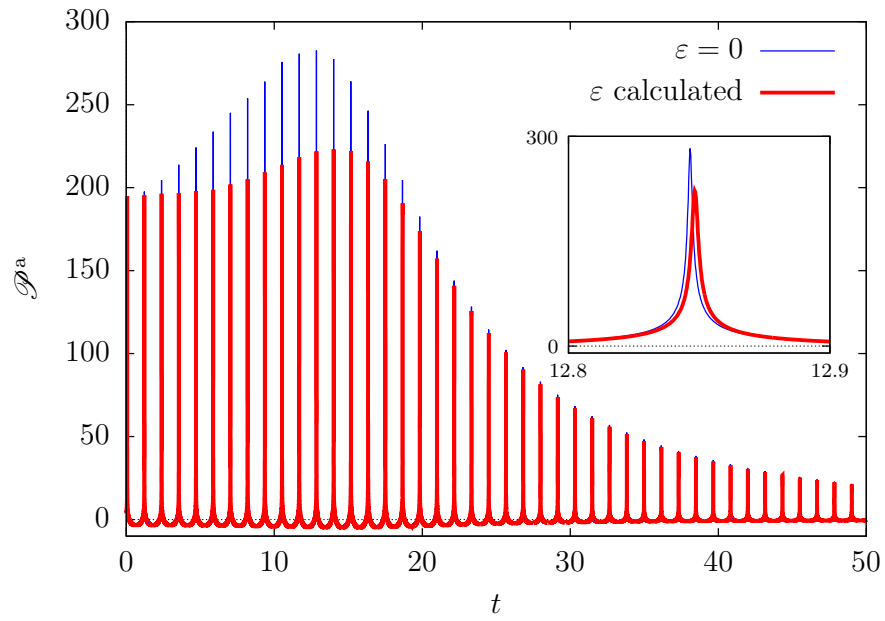
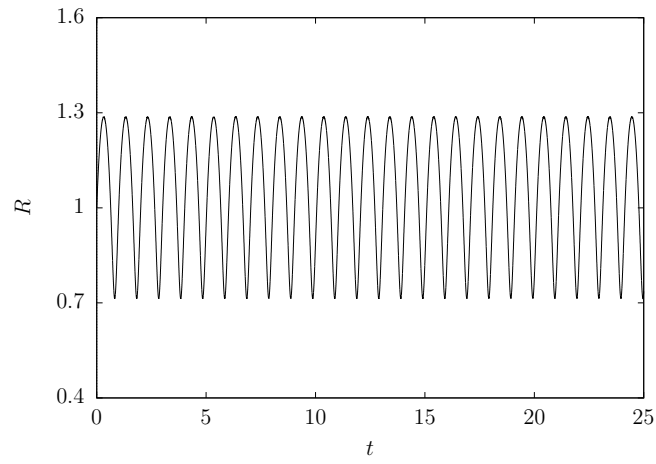
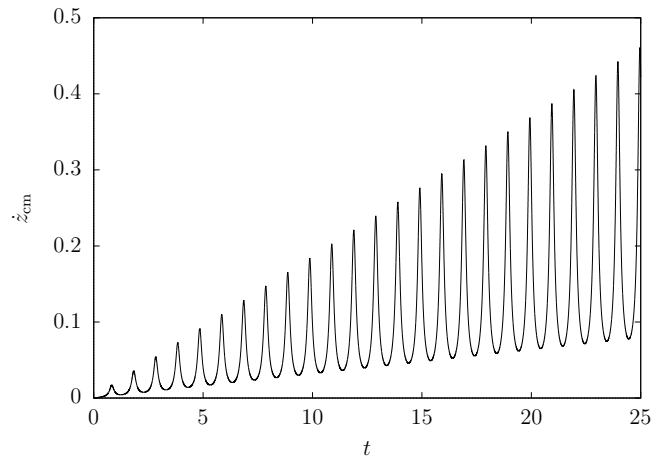


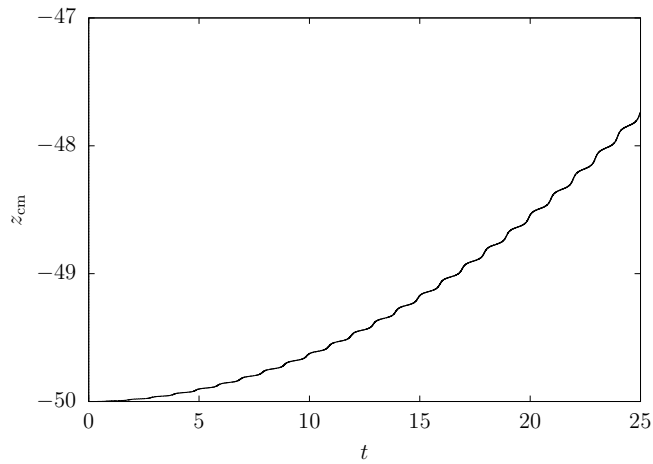
Figure 3: (Color online) Pressure disturbance, and zoom view of a relevant peak for the bubble with vertical motion represented in Fig. 2. The observer is at a distance $r_o = 0.01$ m from the bubble path.



(a)



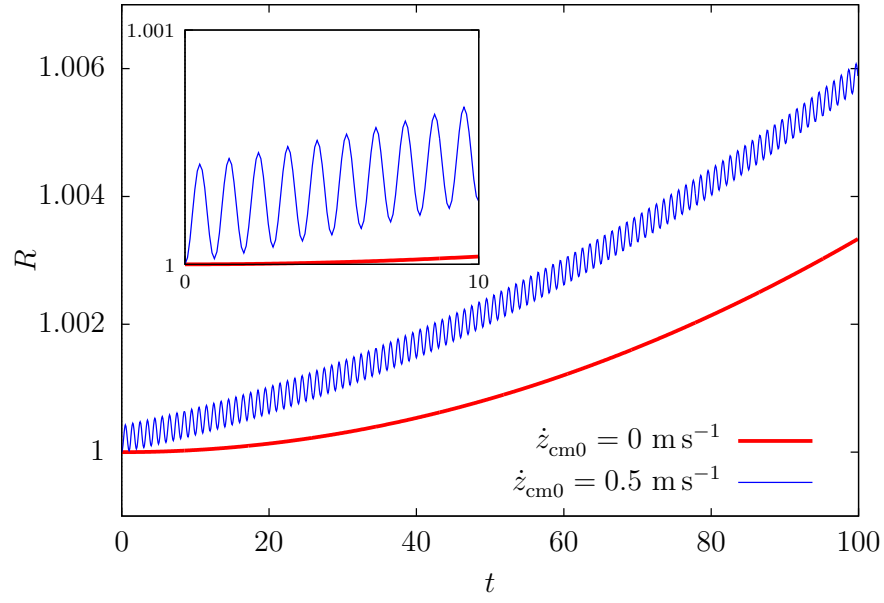
(b)



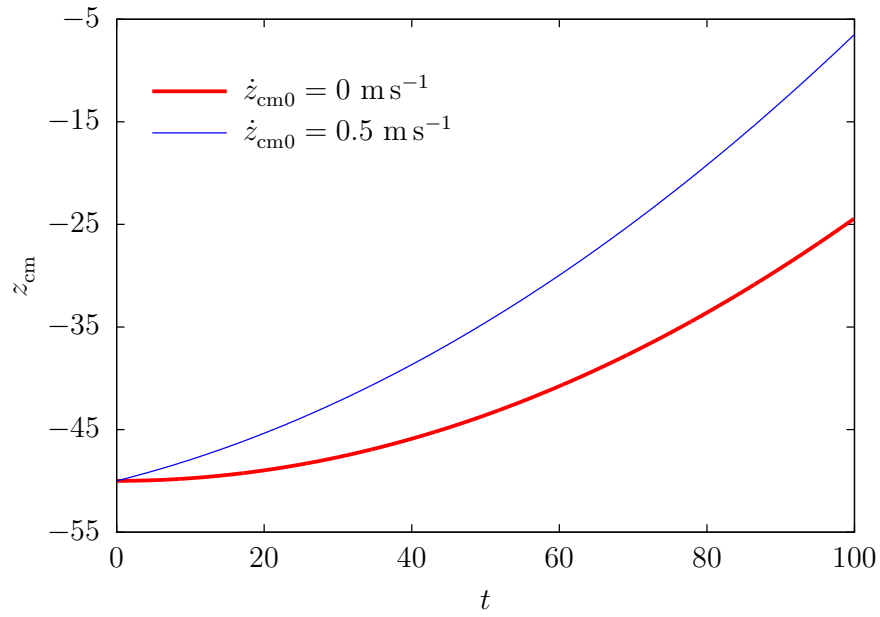
(c)

Figure 4: Bubble radius (a), center of mass velocity (b) and position (c). Initial data:

$$\dot{R}_0 = 5 \text{ m s}^{-1}, \dot{z}_{\text{cm}0} = 0 \text{ m s}^{-1}.$$



(a)



(b)

Figure 5: (Color online) Bubble radius (a) and center of mass position (b). Initial data:

$$\dot{R}_0 = 0 \text{ m s}^{-1}, \dot{z}_{\text{cm}0} = 0.5 \text{ m s}^{-1} \text{ and } \dot{z}_{\text{cm}0} = 0 \text{ m s}^{-1}.$$

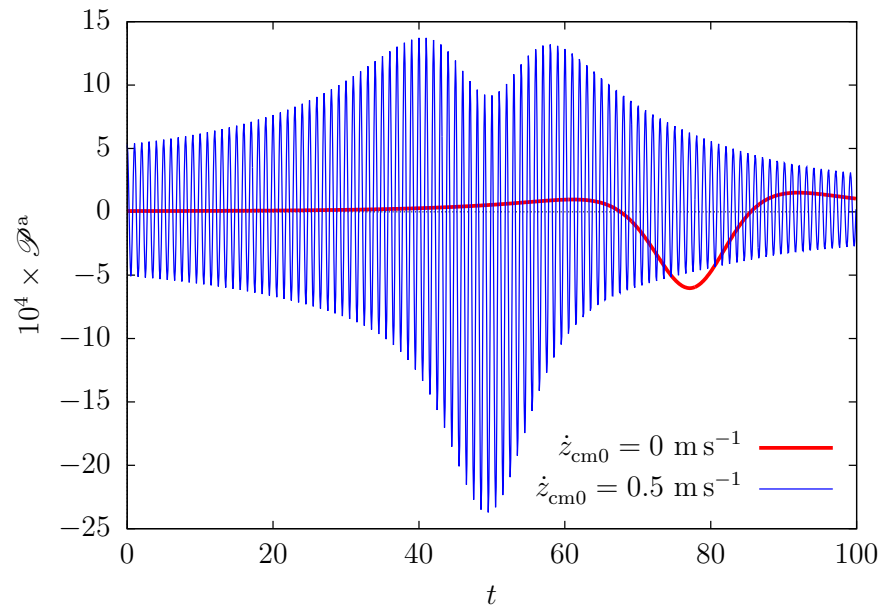


Figure 6: (Color online) Pressure disturbance for the bubble whose motions are shown in Fig. 5. The observer is at a distance $r_o = 0.01$ m from the bubble path.

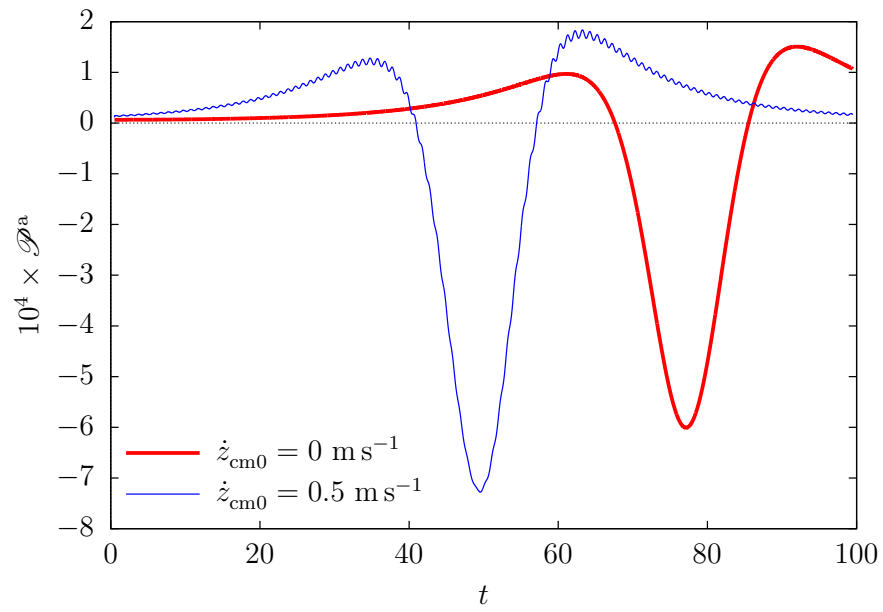


Figure 7: (Color online) Average curves of the pressure disturbances shown in Fig. 6.

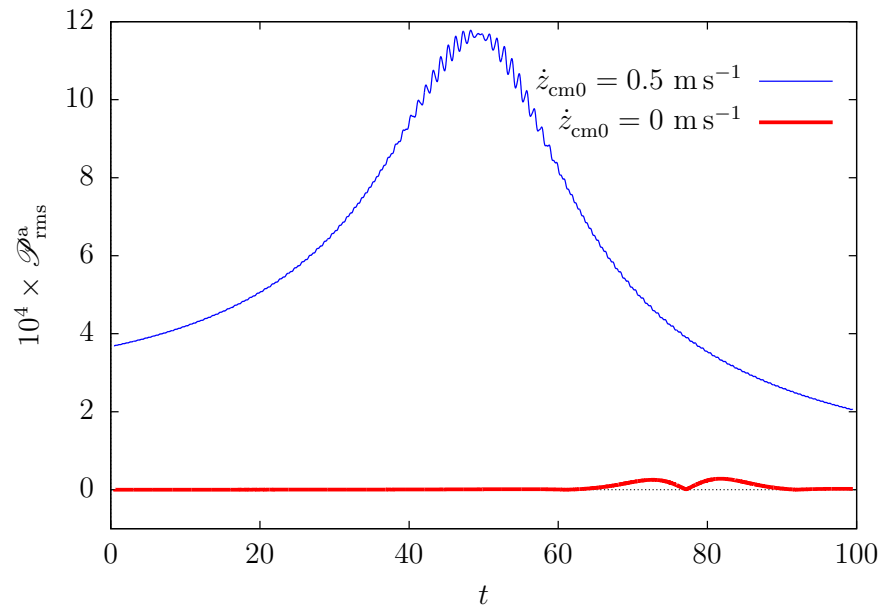


Figure 8: (Color online) Root mean square curves of the pressure disturbances shown in Fig. 6.

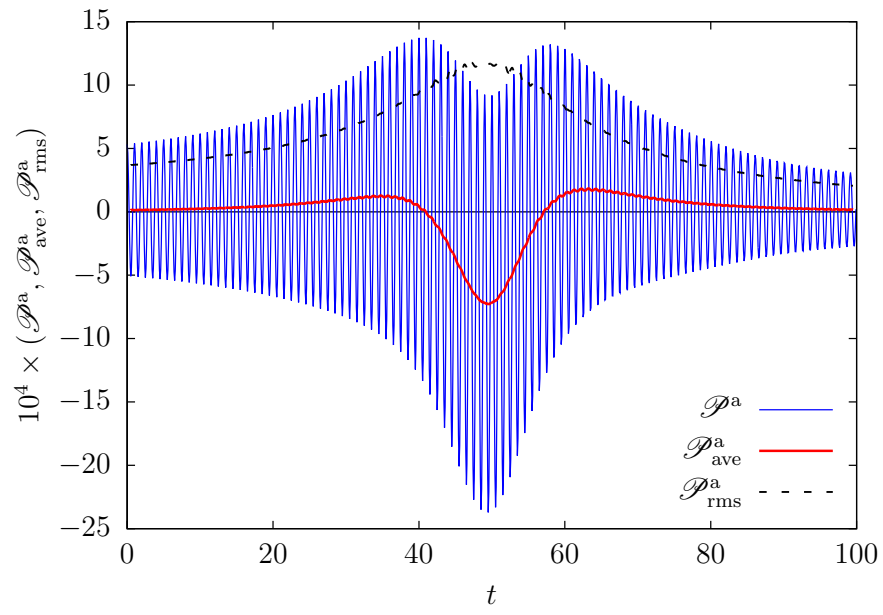


Figure 9: (Color online) Pressure disturbances shown in Fig. 6, and corresponding average and root mean square curves.

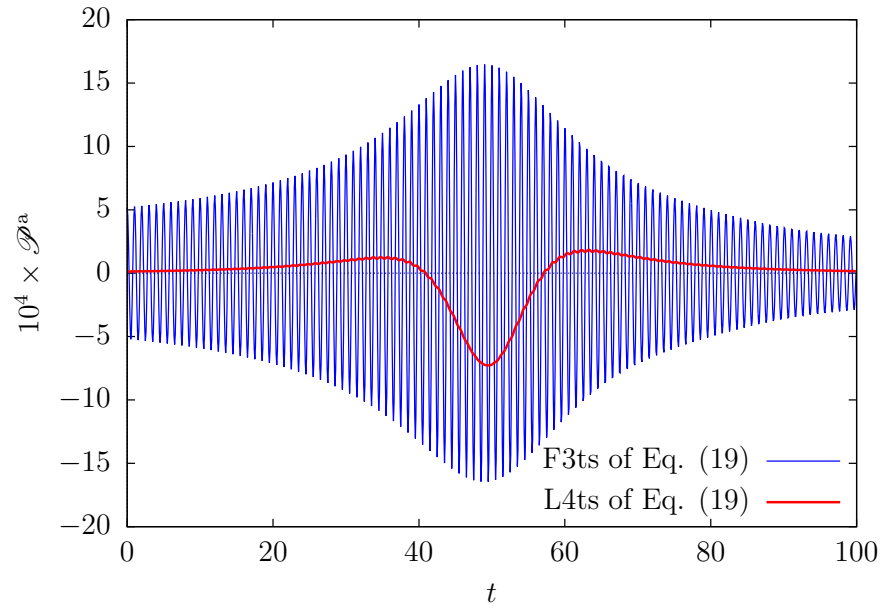


Figure 10: (Color online) Curve of the pressure disturbances, for the bubble whose motions are shown in Fig. 5 (with $\dot{z}_{cm0} = 0.5 \text{ m s}^{-1}$), calculated using the first three terms (F3ts) or the last two terms (L4ts) of equation (19).

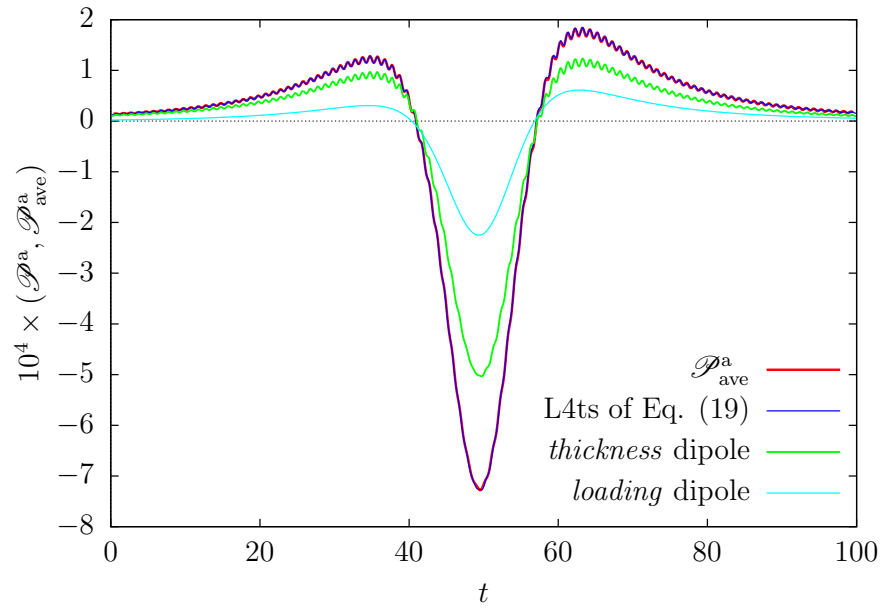


Figure 11: (Color online) Pressure disturbance calculated using the last four terms (L4ts) of equation (19) and average curve of the overall signature, for the bubble whose motions are shown in Fig. 5 (with $\dot{z}_{\text{cm}0} = 0.5 \text{ m s}^{-1}$). The former is the sum of *thickness* and *loading* dipole terms. These contributions are also plotted.

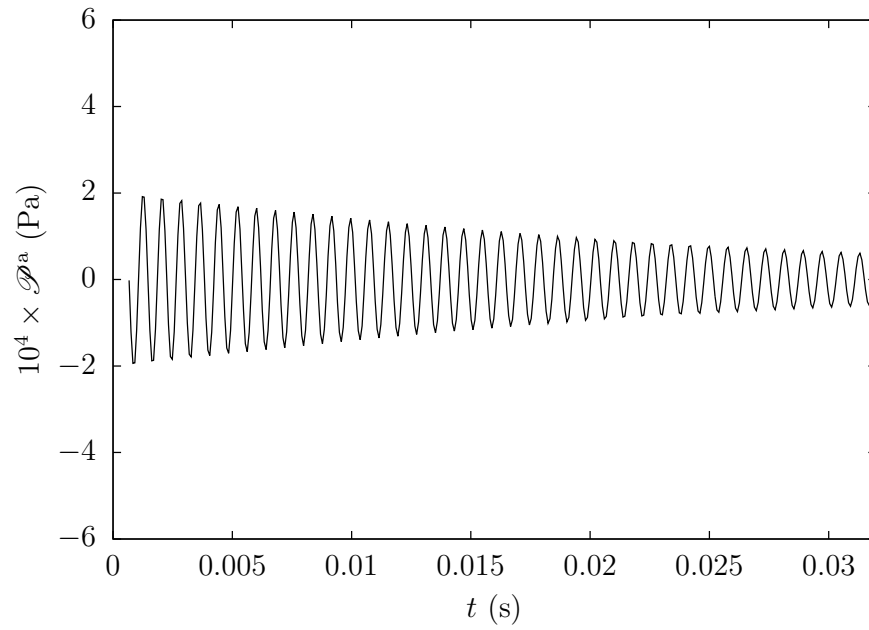


Figure 12: Far-field acoustic signature calculated using thickness noise only to compare with experimental acoustic signature for a single rising bubble reported in Strasberg (1956). Quantities in this figure are given in dimensional form.